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Centre number 

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Surname Matheson

Forename(s) Lewis

Candidate signature 

I declare this is my own work.

# A-level PHYSICS

Paper 3  
Section B Astrophysics

*A Level Physics Online . com*

## Materials

For this paper you must have:

- a pencil and a ruler
- a scientific calculator
- a Data and Formulae Booklet
- a protractor.

## Instructions

- Use black ink or black ball-point pen.
- Fill in the boxes at the top of this page.
- Answer **all** questions.
- You must answer the questions in the spaces provided. Do not write outside the box around each page or on blank pages.
- If you need extra space for your answer(s), use the lined pages at the end of this book. Write the question number against your answer(s).
- Do all rough work in this book. Cross through any work you do not want to be marked.
- Show all your working.

## Information

- The marks for questions are shown in brackets.
- The maximum mark for this paper is 35.
- You are expected to use a scientific calculator where appropriate.
- A Data and Formulae Booklet is provided as a loose insert.

Time allowed: The total time for both sections of this paper is 2 hours. You are advised to spend approximately 50 minutes on this section.

For Examiner's Use	
Question	Mark
1	
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<b>TOTAL</b>	



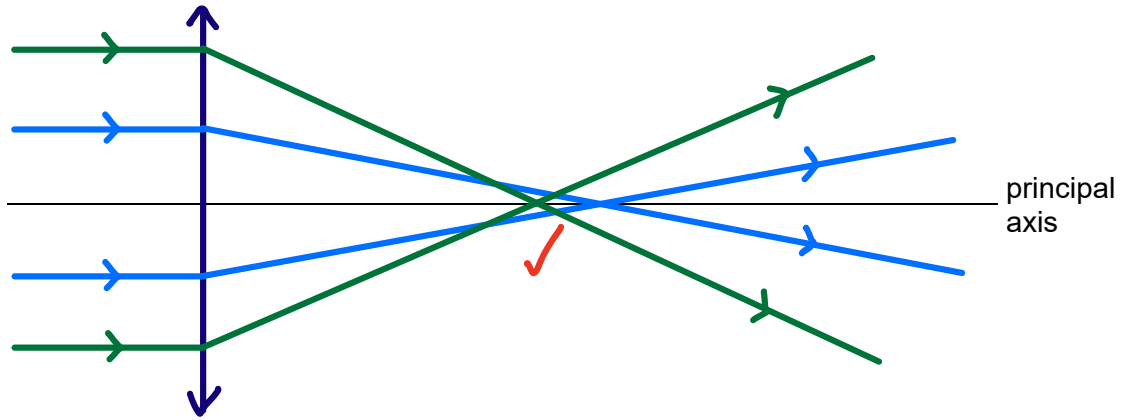
**Section B**

Answer **all** questions in this section.

0 1 . 1

Draw a ray diagram to show how a converging lens can cause spherical aberration.

[1 mark]

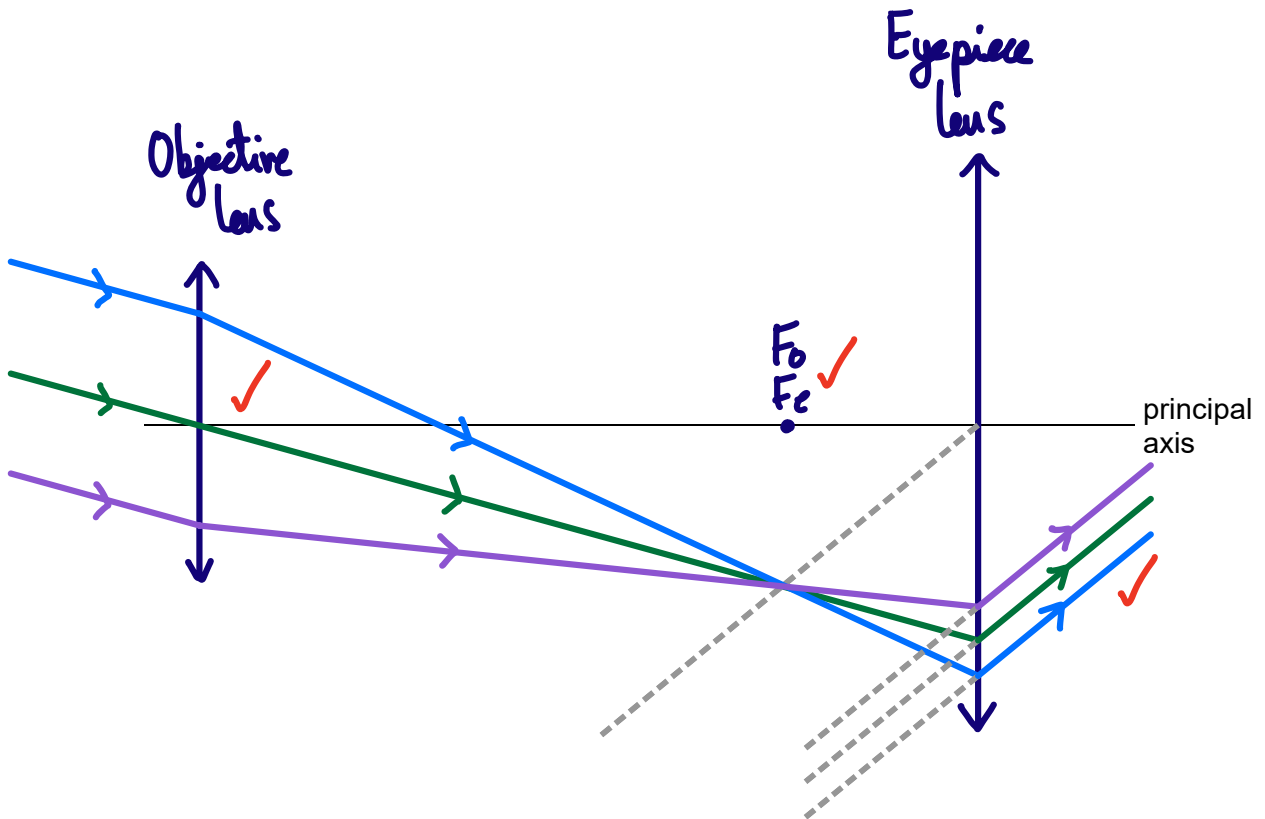


0 1 . 2

Draw a labelled ray diagram for an astronomical refracting telescope in normal adjustment.

Show **three** non-axial rays passing through both lenses. Label the principal foci of the lenses.

[3 marks]



0 1 . 3

The James Lick telescope is an astronomical refracting telescope. When in normal adjustment, the distance between the lenses of the telescope is 17.4 m and the angular magnification is 750

Calculate the focal length of the eyepiece lens.

[2 marks]

$$M = \frac{f_o}{f_e} = 750$$

$$f_o + f_e = 17.4$$

$$f_o = 17.4 - f_e$$

$$\frac{17.4 - f_e}{f_e} = 750$$

$$\frac{17.4}{f_e} = 751$$

$$f_e = 17.4 / 751$$

$$f_e = 2.317 \times 10^{-3}$$

$$\frac{17.4}{f_e} - 1 = 750 \checkmark$$

focal length =  $2.32 \times 10^{-3}$  m  $\checkmark$

0 1 . 4

The James Lick telescope can be used to identify binary stars.

Two techniques are available using this telescope:

- using a processed image from a CCD, and
- direct observation using the naked eye.

Compare the use of a CCD with the use of the naked eye to observe binary stars with this telescope.

[3 marks]

Resolution of CCD is better, this is due to the high number of pixels  $\checkmark$  so stars can more easily be seen as separate.  $\checkmark$

CCDs have a greater quantum efficiency and can be exposed for a long time, so dimmer stars can be observed.  $\checkmark$



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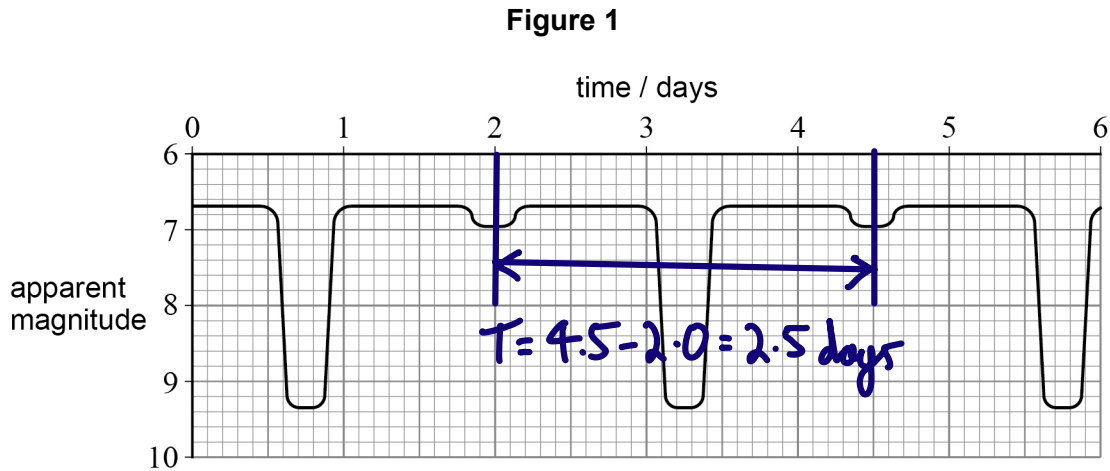


0 2

U Cephei is an eclipsing binary system consisting of two stars that orbit their common centre of mass.

The primary star is class B; the secondary star is class G.

**Figure 1** shows the variation of apparent magnitude of U Cephei with time as observed from Earth.



0 2 . 1

Explain the shape of the graph in **Figure 1**.

[2 marks]

Minima caused by one star passing in front of the other. ✓ The deeper minima is caused by the cooler star (G) passing in front of the brighter star (B). ✓

Question 2 continues on the next page

Turn over ►



A particular spectral line has a wavelength of 486.136 nm when measured from a source in the laboratory.

This line is also present in the absorption spectrum of the primary star of U Cephei. When observed from Earth, the wavelength of the primary star's absorption line varies as shown in **Table 1**.

**Table 1**

	Wavelength / nm
maximum value	486.498
minimum value	485.672

$$\text{Average} = 486.085 < 486.136$$

0 2 . 2 State why the average of the values in **Table 1** is different from the laboratory value.

[1 mark]

The light is blue shifted  $\therefore$  U Cephei is moving towards us. ✓

0 2 . 3 Show that the orbital speed of the primary star is about 250 km s<sup>-1</sup>.

[3 marks]

$$z = \frac{v}{c} = \frac{\Delta\lambda}{\lambda}$$

$$v = \frac{c \Delta\lambda}{\lambda} = \frac{3.00 \times 10^8 \times \left( \frac{486.498 - 485.672}{2} \right)}{486.085} \quad \checkmark$$

$$v = \underline{2.55 \times 10^5 \text{ m s}^{-1}} \checkmark \approx 250 \text{ km s}^{-1}$$



0 2 . 4 Calculate the orbital radius of the primary star.

[2 marks]

$$v = \frac{s}{t} = \frac{2\pi r}{T} \quad r = \frac{vT}{2\pi} = \frac{2.55 \times 10^5 \times 2.5 \times 24 \times 60^2}{2\pi}$$

$$T = 2.5 \text{ days from Fig 1} \checkmark$$

$$r = 8.76 \times 10^9$$

orbital radius =  $8.8 \times 10^9$   $\checkmark$  m

0 2 . 5 Which absorption lines would be most prominent in the spectrum of the primary star?  
Tick ( $\checkmark$ ) **one** box.

[1 mark]

hydrogen

hydrogen and helium

$\checkmark$

ionised metals

neutral metals

0 2 . 6 A different eclipsing binary star system is thought to consist of a white dwarf star and a neutron star.

Discuss how astronomers could confirm this.

[2 marks]

White dwarf in the O or B spectral class  
and has a high temperature, but it is not bright.  $\checkmark$   
Neutron stars have periodic radio emissions  
which would be blocked by white dwarf  
when it passes between us and the neutron  
star.  $\checkmark$



0 3

3C 273 was the first quasar to be discovered.  
IC 1101 is one of the largest galaxies known.  
**Table 2** shows some information about these objects.

Table 2

	Absolute magnitude	Apparent magnitude	Distance / Mpc
quasar 3C 273	X	12.8	760
galaxy IC 1101	-22.8	14.7	320

0 3 . 1

State the property of the quasar that led to its discovery.

[1 mark]

Very powerful radio emissions. ✓

0 3 . 2

Show that the absolute magnitude X of quasar 3C 273 is about -27

[2 marks]

$$m - M = 5 \log \left( \frac{d}{10} \right) \checkmark$$

$$M = m - 5 \log \left( \frac{d}{10} \right) = 12.8 - 5 \log \left( \frac{760 \times 10^6}{10} \right)$$

$$M = \underline{-26.6} \checkmark \approx -27$$





0 3 . 3

Assume that the quasar and the galaxy are both viewed from the same distance.

Explain which would be the brighter object.

Go on to calculate the ratio  $\frac{\text{brightness of brighter object}}{\text{brightness of dimmer object}}$ .

$$M_{\text{Quasar}} = -26.6 \quad M_{\text{Galaxy}} = -22.8 \quad [3 \text{ marks}]$$

The magnitude of the quasar is more negative so it is the brighter object. ✓

Difference of 1 equal to 2.51 times brighter

$$\frac{I_2}{I_1} = 2.51^{M_1 - M_2} = 2.51^{(-22.8 - (-26.6))}$$

$$= 2.51^{3.8} = 33 \quad \checkmark$$

$$\text{ratio} = \underline{33} \quad \checkmark$$

0 3 . 4

The black hole at the centre of IC 1101 has a mass of  $7.1 \times 10^{11} M_{\text{S}}$  where  $M_{\text{S}}$  is the mass of the Sun.

Calculate the average density within the event horizon of the black hole.

$$R_s = \frac{2GM}{c^2} = \frac{2 \times 6.67 \times 10^{-11} \times 7.1 \times 10^{11} \times 1.99 \times 10^{30}}{(3.00 \times 10^8)^2} \quad [3 \text{ marks}]$$

$$R_s = 2.094 \times 10^{15} \text{ m} \quad \checkmark$$

$$V = \frac{4}{3} \pi R_s^3 = \frac{4}{3} \pi \times (2.094 \times 10^{15})^3 = 3.847 \times 10^{46} \text{ m}^3 \quad \checkmark$$

$$\rho = \frac{m}{V} = \frac{7.1 \times 10^{11} \times 1.99 \times 10^{30}}{3.847 \times 10^{46}} = 3.672 \times 10^{-5}$$

$$\text{average density} = \underline{3.7 \times 10^{-5}} \quad \checkmark$$

kg m<sup>-3</sup>

9

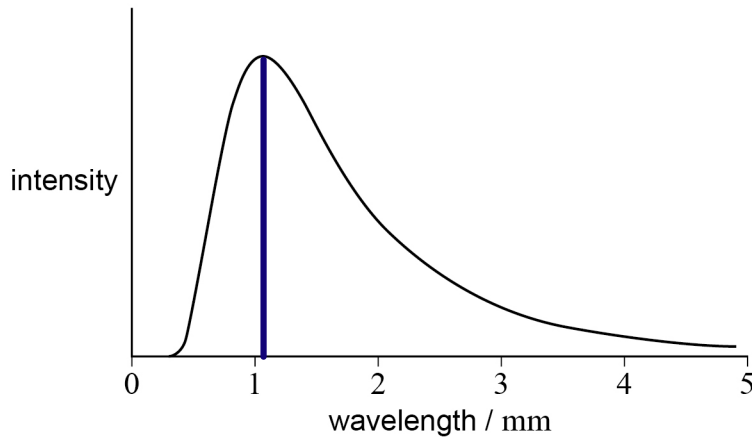
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0 4

In the middle of the 20th century, there were two competing theories of the Universe. In 1964, electromagnetic radiation was observed coming from all directions in space. **Figure 2** shows the distribution of this radiation as observed from Earth.

Figure 2



The graph provides evidence for one of these theories of the Universe.

Discuss the main features of this theory of the Universe. → Big Bang

In your answer, you should include:

- the main predictions and evidence for the theory, and
- a suitable calculation.

[6 marks]

Big Bang: universe has expanded from a single point.

Redshift: Theory states all distant galaxies are moving away from us with the recession velocity increasing with distance due to expansion of universe. ✓✓



CMBR: Theory predicts black body radiation at microwave wavelengths from all directions, indicating the universe used to be small, hot and dense but has since expanded. ✓✓

The peak of graph in microwave region  $\approx 1.1 \times 10^{-3}$  m and has a curve consistent with black body radiation.

$$\text{Wien's law } \lambda_{\text{peak}} T = 2.9 \times 10^{-3}$$

$$T = \frac{2.9 \times 10^{-3}}{1.1 \times 10^{-3}}$$

$$T = 2.64 \text{ K}$$

$\therefore$  consistent with theory. ✓✓

END OF QUESTIONS



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